# Markscheme 

May 2018

Physics

## Standard level

## Paper 3

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Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1. | a |  | smooth line, not kinked, passing through all the error bars $\checkmark$ |  | 1 |
| 1. | b | i | $0.84 \pm 0.03$ « ${ }^{\text {» }}$ Ј | Accept any value from the range: 0.81 to 0.87 . <br> Accept uncertainty 0.03 OR 0.025. | 1 |
| 1. | b | ii | $\begin{aligned} & K=\sqrt{0.005} \times 0.84=0.059 \\ & \text { « } \frac{\Delta K}{K}=\frac{\Delta P}{P} » \\ & \Delta K=\frac{0.03}{0.84} \times 0.0594=0.002 \\ & « K=(0.059 \pm 0.002) » \end{aligned}$ <br> uncertainty given to 1 sf $\checkmark$ | Allow ECF [3 max] if 10T is used. Award [3] for BCA. | 3 |
| 1. | b | iii | $s T^{\frac{1}{2}}$ | Accept $s \sqrt{T}$ or in words. | 1 |
| 1. | c |  | straight AND ascending line $\checkmark$ through origin |  | 2 |
| 1. | d |  | $K=\sqrt{\text { slope }} \checkmark$ |  | 1 |


| 2. | a |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |

## Section B

## Option A — Relativity

| 3. | $\mathbf{a}$ |  | magnetic field $\checkmark$ |  |
| :--- | :--- | :--- | :--- | :--- |
| 3. | $\mathbf{b}$ | $\mathbf{i}$ | «according to Y» the positive charges are moving «to the right» $\checkmark$ <br> $d$ decreases $\checkmark$ | For MP1, movement of positive charges <br> must be mentioned explicitly. |
| 3. | $\mathbf{b}$ | ii | 2 <br> positive charges are moving, so there is a magnetic field $\checkmark$ <br> the density of positive charges is higher than that of negative charges, <br> so there is an electric field $\checkmark$ | The reason must be given for each <br> point to be awarded. |


| 4. | a | i | $\text { « } \frac{10^{4}}{0.995 \times 3 \times 10^{8}}=» 34<\mu s » \checkmark$ | Do not accept $10^{4} / \mathrm{c}=33 \mu \mathrm{~s}$. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4. | a | ii | time is much longer than 10 times the average life time «so only a small proportion would not decay" $\checkmark$ |  | 1 |
| 4. | b | i | $\begin{aligned} & \gamma=10 \checkmark \\ & \Delta t_{0}=« \frac{\Delta t}{\gamma}=\frac{34}{10}=» 3.4 « \mu s » \end{aligned}$ |  | 2 |
| 4. | b | ii | the value found in (b)(i) is of similar magnitude to average life time $\checkmark$ significant number of muons are observed on the ground $\checkmark$ «therefore this supports the special theory» |  | 2 |

5. a $^{\prime 2}$
(continued...)
(Question 5 continued)
6. L $^{\text {b }}$

## Option B - Engineering physics

| 6. | a |  | $\begin{aligned} & \Gamma «=F r=50 \times 2 »=100 « \mathrm{Nm} » \checkmark \\ & \alpha «=\frac{\Gamma}{I}=\frac{100}{450} »=0.22 « \mathrm{rads}^{-2} » \checkmark \end{aligned}$ | Final value to at least 2 sig figs, $\mathbf{O R}$ clear working with substitution required for mark. | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 6. | b | i | $\begin{aligned} & « \omega_{t}^{2}-\omega_{0}^{2}=2 \alpha \Delta \theta » \\ & « \omega_{t}^{2}-0=2 \times 0.22 \times 2 \pi » \\ & \omega_{t}=1.7 \text { «rads }^{-1} » \checkmark \end{aligned}$ | Accept BCA, values in the range: 1.57 to 1.70 . | 1 |
| 6. | b | ii | $\begin{aligned} & « L=I \omega=450 \times 1.66 » \\ & =750<\mathrm{kgm}^{2} \mathrm{rads}^{-1} » \end{aligned}$ | Accept BCA, values in the range: 710 to 780. | 1 |
| 6. | c |  | $\begin{aligned} & « I=450+m r^{2} » \\ & I «=450+30 \times 2^{2} »=570 \text { « } \mathrm{kgm}^{2} » \checkmark \\ & « L=570 \times \omega=747 » \\ & \omega=1.3 \text { «rads}{ }^{-1} » \checkmark \end{aligned}$ | Watch for ECF from (a) and (b). <br> Accept BCA, values in the range: 1.25 to 1.35 . | 2 |

(continued...)
(Question 6 continued)

| 6. | $\mathbf{d}$ | $\mathbf{i}$ | moment of inertia will decrease $\checkmark$ <br> angular momentum will be constant «as the system is isolated» $\checkmark$ <br> «so the angular speed will increase» |  |
| :--- | :--- | :--- | :--- | :--- |
| 6. | $\mathbf{d}$ | $\mathbf{i i}$ | $\omega_{t}=1.66$ from bi $A N D W=\Delta E_{k} \checkmark$ <br> $W=\frac{1}{2} \times 450 \times 1.66^{2}-\frac{1}{2} \times 570 \times 1.31^{2}=131 « J » \checkmark$ | ECF from 8bi |
| Accept BCA, value depends on the |  |  |  |  |
| answers in previous questions. |  |  |  |  |


| 7. | $\mathbf{a}$ | " $p_{1} V_{1}^{\frac{5}{3}}=p_{2} V_{2}^{\frac{5}{3}} »$ <br> $1.1 \times 10^{5} \times 5^{\frac{5}{3}}=p_{2} \times 2^{\frac{5}{3}} \checkmark$ <br> $p_{2} «=\frac{1.1 \times 10^{5} \times 5^{\frac{5}{3}}}{2.5^{\frac{5}{3}}} »=5.066 \times 10^{5} « \mathrm{~Pa} » \checkmark$ | Volume may be in litres or $m^{3}$. <br> Value to at least 2 sig figs, OR clear working with <br> substitution required for mark. |
| :--- | :--- | :--- | :--- |

(Question 7 continued)


## Option C - Imaging

| 8. | a | i | image is real «as projected on a screen» $\checkmark$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8. | a | ii | $\begin{aligned} & «-\frac{18}{u}=-0.40 » \\ & u=45 \\ & \frac{1}{45}+\frac{1}{18}=\frac{1}{f} \end{aligned}$ <br> OR $\begin{aligned} & f=13 « \mathrm{~cm} » \\ & P=\frac{1}{f}=« \frac{1}{13} »=0.078<\mathrm{cm}^{-1} » \end{aligned}$ | Accept answer 7.7«D». | 3 |
| 8. | a | iii | refractive index depends on wavelength $\checkmark$ light of different wavelengths have different focal points / refract differently $\checkmark$ there will be coloured fringes around the image / image will be blurred $\checkmark$ |  | 3 |
| 8. | b |  | any 2 correct rays to find image from lens $1 \checkmark$ ray to locate $F_{2} \checkmark$ <br> focal length $=$ « $-» 70$ «cm» $\checkmark$ | Accept values in the range: 65 cm to 75 cm . <br> Accept correct MP3 from accepted range also if working is incorrect or unclear, award [1]. | 3 |

(continued...)
(Question 8 continued)


| 9. | a |  | $\begin{aligned} & « \sin c=\frac{1.34}{1.56} 》 \\ & c=59.2 «^{\circ} » \end{aligned}$ | Accept values in the range: 59.0 to 59.5 . Accept answer 1.0 rad. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9. | b |  | optic fibres are not susceptible to earthing problems $\checkmark$ optic fibres are very thin and so do not require the physical space of electrical cables $\checkmark$ <br> optic fibres offer greater security as the lines cannot be tapped $\checkmark$ optic fibres are not affected by external electric/magnetic fields/interference $\checkmark$ <br> optic fibres have lower attenuation than electrical conductors / require less energy $\sqrt{ }$ <br> the bandwidth of an optic fibre is large and so it can carry many communications at once/in a shorter time interval /faster data transfer $\checkmark$ |  | 2 max |
| 9. | c | i | a signal that is wider and lower, not necessarily rectangular, but not a larger area $\checkmark$ |  | 1 |
| 9. | c | ii | $\begin{aligned} & \text { attenuation }=-1.24 \times 3.4 «=-4.216 \mathrm{~dB} » \\ & -4.216=10 \log \frac{I}{15} \checkmark \\ & I=5.68 « \mathrm{~mW} » \end{aligned}$ | Need negative attenuation for MP1, may be shown in MP2. <br> For mp3 answer must be less than 15 mW (even with ECF) to earn mark <br> Allow [3] for BCA. | 3 |

(continued...)
(Question 9 continued)

| 9. | c iii | refractive index near the edge of the core is less than at the centre $\checkmark$ <br> speed of rays which are reflected from the cladding are greater than the speed of <br> rays which travel along the centre of the core $\checkmark$ |
| :--- | :--- | :--- | :--- | :--- | :--- |
| the time difference for the rays that reflect from the cladding layer compared to |  |  |
| those that travel along the centre of the core is less |  |  |
| OR |  |  |
| the signal will remain more compact/be less spread out /dispersion is lower $\checkmark$ |  |  |
| bit rate of the system may be greater $\checkmark$ |  |  |$~ 3$ max

## Option D - Astrophysics

| 10. | a | i | a galaxy is much larger in size than a solar system $\checkmark$ <br> a galaxy contains more than one star system / solar system $\checkmark$ <br> a galaxy is more luminous $\checkmark$ | Any other valid statement. | 1 max |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 10. | a | ii | a comet is a small icy body whereas a planet is mostly made of rock or gas $\checkmark$ <br> a comet is often accompanied by a tail/coma whereas a planet is not $\checkmark$ <br> comets (generally) have larger orbits than planets $\checkmark$ <br> a planet must have cleared other objects out of the way in its orbital neighbourhood $\checkmark$ | $\mathbf{1 m m a x}$ |  |


| 11. | a | i | the wavelengths of the dips correspond to the wavelength in the <br> emission spectrum $\checkmark$ <br> the absorption lines in the spectrum of star $X$ suggest it contains predominantly <br> hydrogen <br> OR <br> main sequence stars are rich in hydrogen $\checkmark$ | 2 |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 11. | a | ii | peak wavelength: $290 \pm 10$ «nm» $\checkmark$ <br> $T=\frac{2.9 \times 10^{-3}}{290 \times 10^{-9}}=« 10000 \pm 400 \mathrm{~K} »$ | Substitution in equation must be seen. <br> Allow ECF from MP1. | $\mathbf{2}$ |

(continued...)
(Question 11 continued)

| 11. | b | i | $35 \pm 5 L_{\text {s }} \checkmark$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 11. | b | ii | $\begin{aligned} & \frac{L_{\mathrm{x}}}{L_{\mathrm{s}}}=\frac{R_{\mathrm{x}}^{2} \times \mathrm{T}_{\mathrm{x}}^{4}}{R_{\mathrm{s}}^{2} \times \mathrm{T}_{\mathrm{s}}^{4}} \\ & \text { OR } \\ & R_{\mathrm{x}}=\sqrt{\frac{L_{\mathrm{x}} \mathrm{~T}_{\mathrm{s}}^{4}}{L_{\mathrm{s}}^{4} T_{\mathrm{x}}^{4}}} \times R_{\mathrm{s}} \\ & R_{\mathrm{x}}=\sqrt{\frac{35 \times 6000^{4}}{10000^{4}}} \times R_{\mathrm{s}} \text { (mark for correct substitution) } \\ & R_{\mathrm{x}}=2.1 R_{\mathrm{s}} \end{aligned}$ | Allow ECF from (b)(i). <br> Accept values in the range: 2.0 to $2.3 R_{s}$. <br> Allow $T_{s}$ in the range: 5500 K to 6500 K . | 3 |
| 11. | b | iii | $\begin{aligned} & M_{\mathrm{x}}=(35)^{\frac{1}{3.5}} M_{\mathrm{s}} \checkmark \\ & M_{\mathrm{x}}=2.8 M_{\mathrm{s}} \checkmark \end{aligned}$ | Allow ECF from (b)(i). <br> Do not accept $M_{X}=(35)^{\frac{1}{3.5}}$ for first marking point. Accept values in the range:2.6 to $2.9 M_{s}$. | 2 |
| 11. | c |  | the star «core» collapses until the «inward and outward» forces / pressures are balanced $\checkmark$ <br> the outward force / pressure is due to electron degeneracy pressure «not radiation pressure» $\checkmark$ |  | 2 |


| 12. | a | experiments and collecting data are extremely costly $\checkmark$ data from many projects around the world can be collated $\checkmark$ | OWTTE | 1 max |
| :---: | :---: | :---: | :---: | :---: |
| 12. | b | $\begin{aligned} & v=« z c=0.19 \times 3 \times 10^{8}=» 5.7 \times 10^{7}<\mathrm{ms}^{-1} » \checkmark \\ & d=« \frac{v}{H_{0}}=\frac{5.7 \times 10^{4}}{70} »=810 \mathrm{Mpc} \text { OR } 8.1 \times 10^{8} \mathrm{pc} \checkmark \end{aligned}$ | Correct units must be present for MP2 to be awarded. <br> Award [2] for BCA. | 2 |
| 12. | C | aLTERNATIVE 1 $\begin{aligned} & \frac{R_{\text {now }}}{R_{\text {then }}}=1+z=1.19 \checkmark \\ & \text { so (assuming constant expansion rate) } \frac{t_{\text {now }}}{t}=1.19 \\ & t=\frac{14}{1.19}=11.7 \mathrm{By}=12 \text { «By (billion years)" } \end{aligned}$ <br> ALTERNATIVE 2 <br> light has travelled a distance: $\left(810 \times 10^{6} \times 3.26=\right) 2.6 \times 10^{9} \mathrm{ly} \checkmark$ so light was emitted: 2.6 billion years ago $\checkmark$ so the universe was 11.4 billion years old $\checkmark$ | MP1 can be awarded if MP2 clearly seen. <br> Accept $2.5 \times 10^{25} \mathrm{~m}$ for mp1. <br> MP1 can be awarded if MP2 clearly seen. | 3 |

